

# STATUS OF THE MINIMAL SUPERSYMMETRIC $SO(10)$ MODEL\*

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\*Borut Bajc, I.D. i Miha Nemevšek, arXiv:0809.1069, **JHEP** 0811:007, 2008.

# OUTLINE

- **MINIMAL  $SO(10)$  MODEL: MAIN FEATURES**
- **LOW SUPERSYMMETRIC SCALE & NEUTRINO MASS**
- **SPLIT SUPERSYMMETRY & NEUTRINO MASS**
- **EXPERIMENTAL SIGNALS**
- **CONCLUSIONS**

# MINIMAL $SO(10)$ — MSG(rand)U(nified)T(heory)\*

**REPRESENTATIONS:**  $16_{Fi}, (i = 1, 2, 3)$   $210_H$   $126_H$   $\overline{126}_H$   $10_H$

**SUPERPOTENTIAL:**  $W_Y = 16_F(Y_{10}10_H + Y_{\overline{126}}\overline{126}_H)16_F$

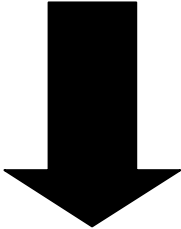
( $W_Y$  is relevant for the description of fermion masses!  
 $Y_{10}$  and  $Y_{\overline{126}}$  are  $3 \times 3$  symmetric Yukawa matrices.)

$$W_H = \frac{m}{4!}210_H^2 + \frac{\lambda}{4!}210_H^3 + \frac{M}{5!}126_H\overline{126}_H + \frac{\eta}{5!}126_H210_H\overline{126}_H + m_H10_H^2 + \frac{1}{4!}210_H10_H(\alpha126_H + \overline{\alpha}\overline{126}_H)$$

( $W_H$  is relevant for the gauge coupling unification ( $\alpha_i, i = 1, 2, 3$ )!)

\*T. E. Clark, T. K. Kuo and N. Nakagawa, Phys. Lett. B 115 (1982) 26; K. S. Babu and R. N. Mohapatra, Phys. Rev. Lett. 70 (1993) 2845; C. S. Aulakh, B. Bajc, A. Melfo, G. Senjanović and F. Vissani, Phys. Lett. B 588 (2004) 196; ...

## RELEVANT PARAMETERS: ( $W_Y$ )

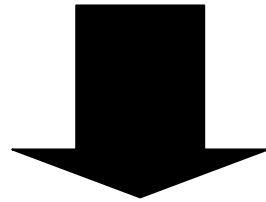
EXPERIMENTALLY DETERMINED QUANTITIES	MODEL
3 masses of “up” quarks: $m_u, m_c, m_t$	$Y_{10}$ $Y_{126}$
3 masses of “down” quarks: $m_d, m_s, m_b$	
3 masses of charged leptons: $m_e, m_\mu, m_\tau$	
2 parameters (3 masses) of neutral leptons: $m_1, m_2, m_3$	
3 CKM* angles + 1 CKM phase: $\theta_{12, 23, 13}^{CKM}, \delta^{CKM}$	
2 (3) PMNS† angles + 0 (3) PMNS phases: $\theta_{12, 23, 13}^{PMNS}, \dots$	$3 + 12$
<b>17 (22)</b>	<b>15</b>

\*Cabibbo-Kobayashi-Maskawa (CKM) (+Nobel Prize in Physics 2008)

†Pontecorvo-Maki-Nakagawa-Sakata (PMNS)

## RELEVANT PARAMETERS: ( $W_H$ )

**CONDITIONS\*:**  $M, m_H, m \longrightarrow M = m \frac{\eta}{\lambda} \frac{3 - 14x + 15x^2 - 8x^3}{(x-1)^2} \longrightarrow m$



### PARAMETERS\*:

$$m, \alpha, \bar{\alpha}, |\lambda|, |\eta|, \phi = \arg(\lambda) = -\arg(\eta), x = \text{Re}(x) + i\text{Im}(x), g_{GUT}, M_{SUSY}$$

**These parameters are important for the running of gauge couplings.**

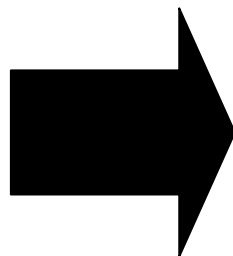
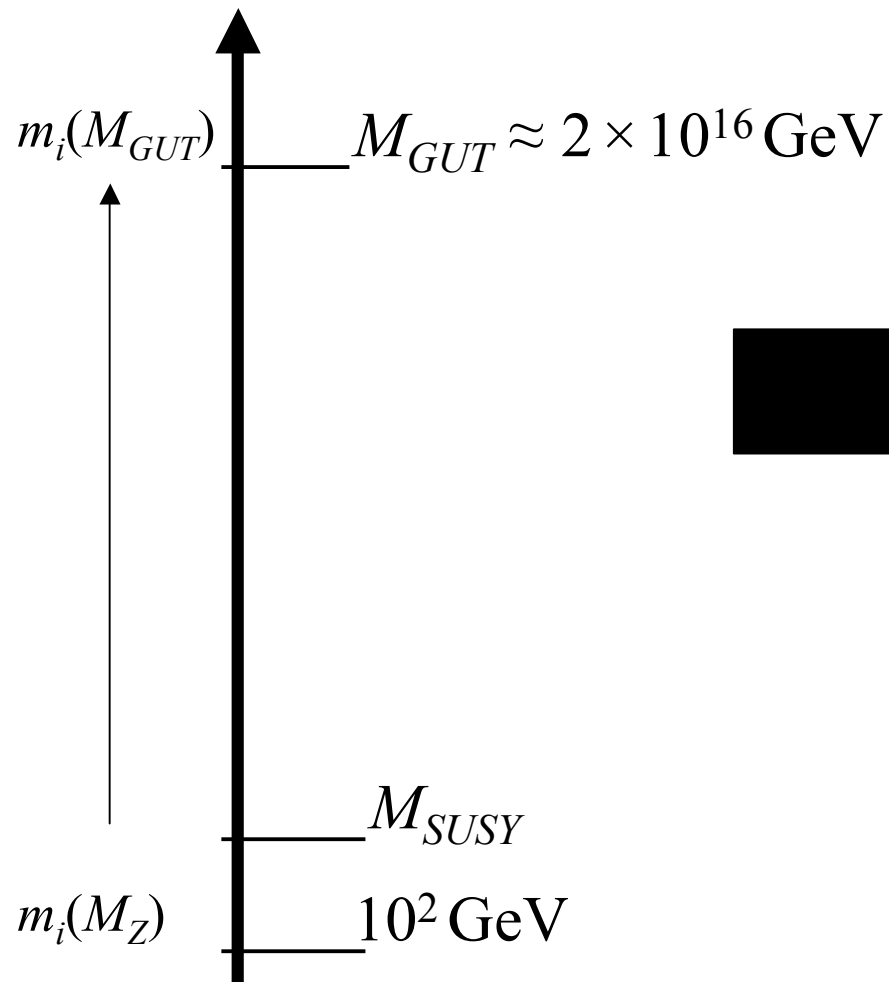
$$\tan \beta$$

**These parameters are important for the running of Yukawa couplings.**

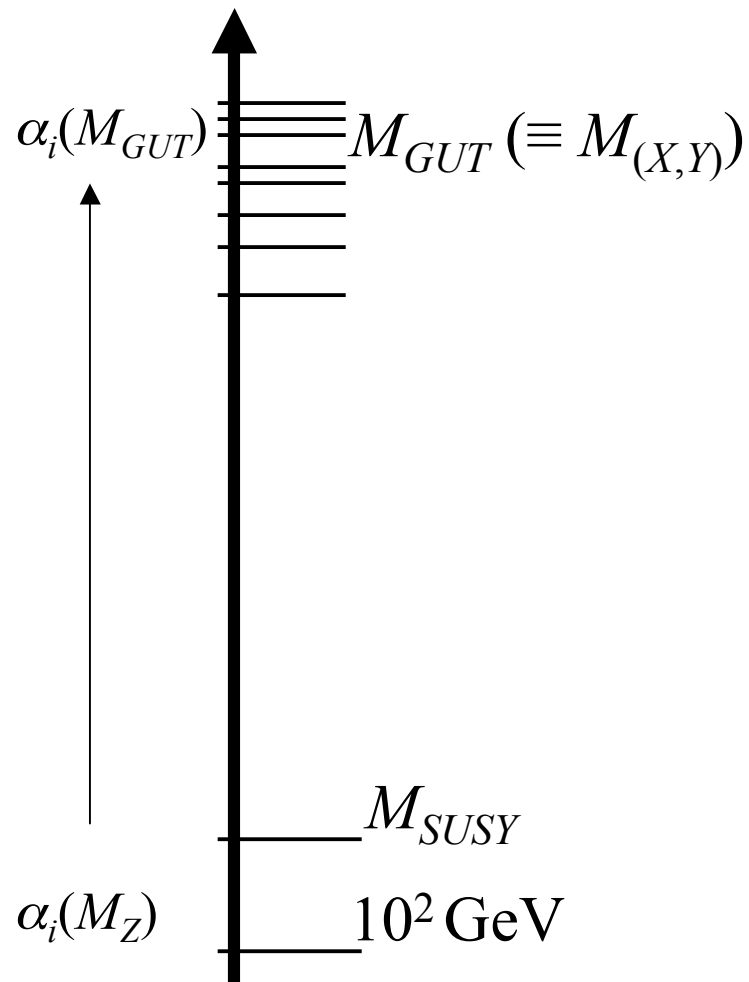
\*C. S. Aulakh, B. Bajc, A. Melfo, G. Senjanović and F. Vissani, Phys. Lett. B 588 (2004) 196.

# UNIFICATION

Minimal Supersymmetric Standard Model (MSSM)



MSGUT

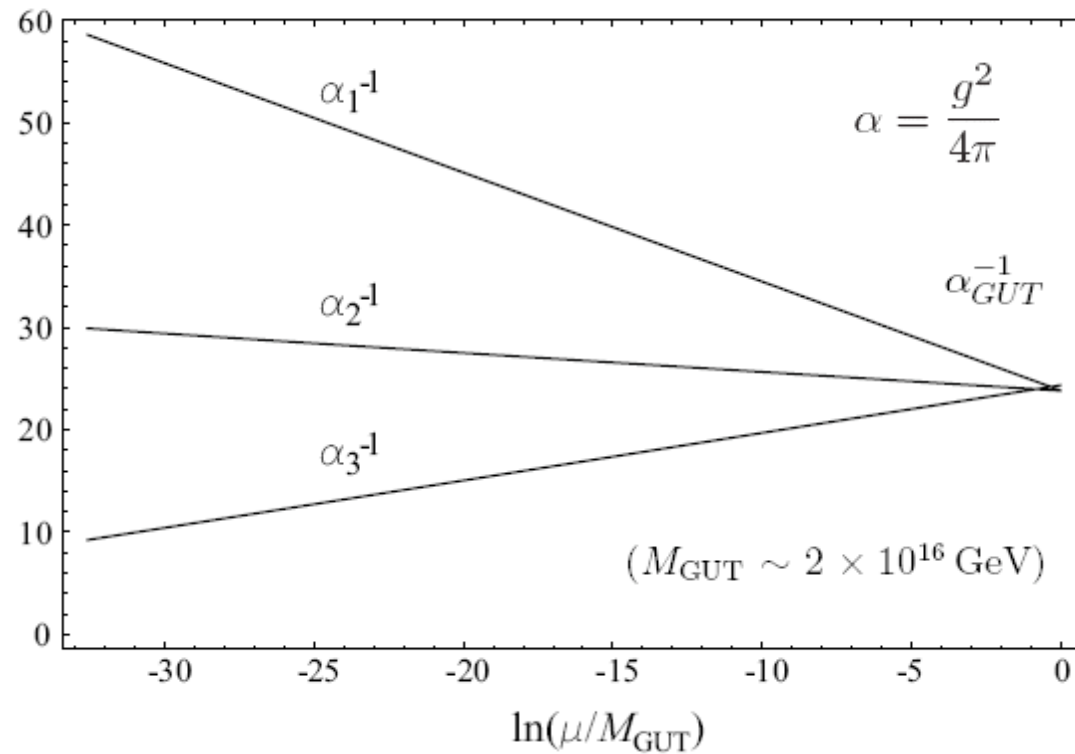


# MSSM: GAUGE COUPLING UNIFICATION

$$\alpha_1(M_Z) = 0.016949 \pm 0.000005$$

$$\alpha_2(M_Z) = 0.033816 \pm 0.000027$$

$$\alpha_3(M_Z) = 0.1176 \pm 0.0020$$



# MSGUT: GAUGE COUPLING UNIFICATION\*



\*B. Bajc et al, hep-ph/0402122; C. S. Aulakh and S. K. Garg, hep-ph/0512224; S. Bertolini, T. Schwetz and M. Malinsky, hep-ph/0605006; ...



## FERMION MASSES IN MINIMAL $SO(10)$ \*

$$M_u = \frac{N_u}{N_d} \tan \beta [M_d + \xi(M_d - M_e)]$$

$$M_n = \frac{v \sin^2 \beta}{m \cos \beta} \alpha \sqrt{\frac{|\lambda|}{|\eta|}} \frac{N_u^2}{N_d} [m_I f_I + m_{II} f_{II}]$$

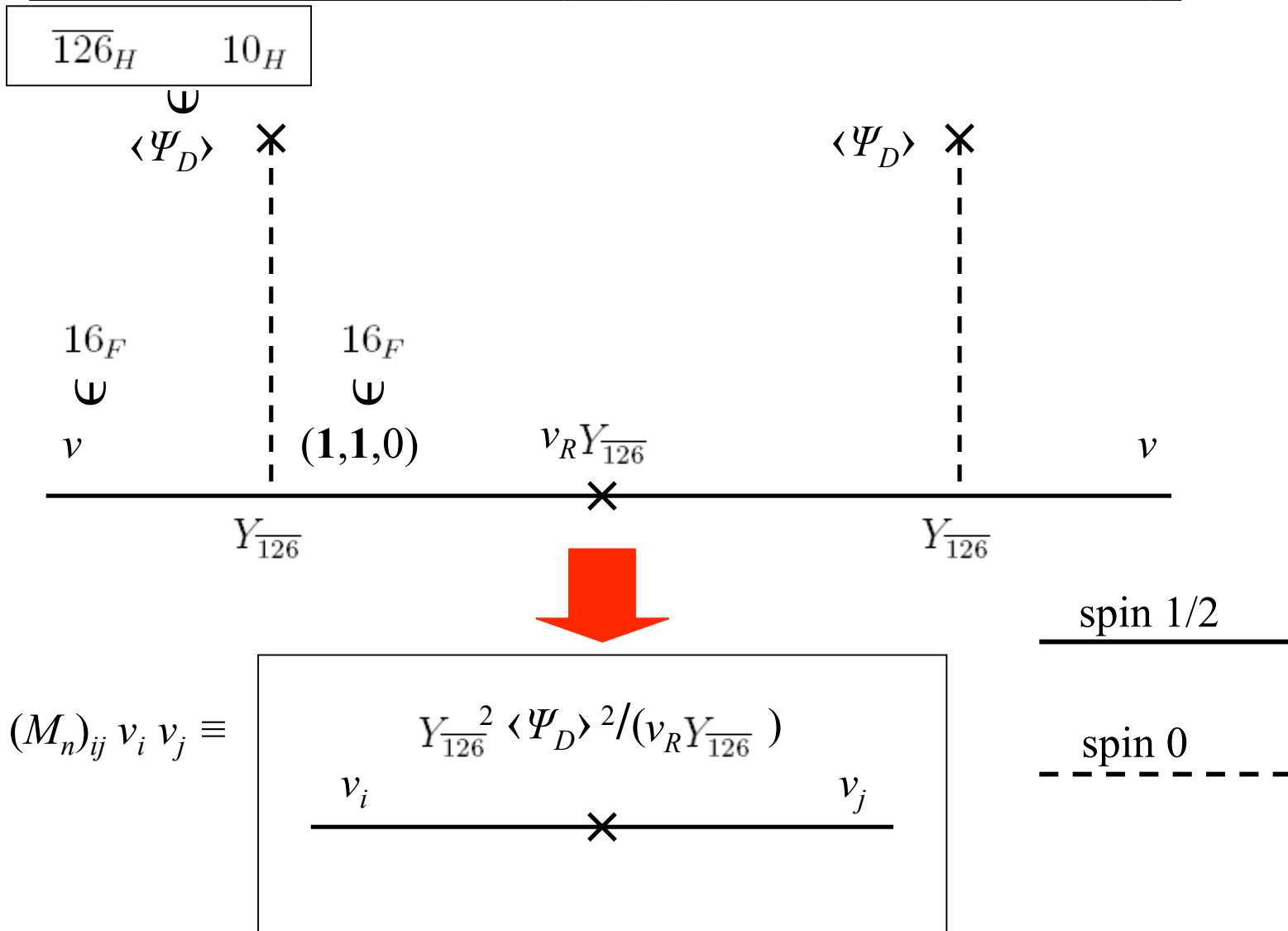
“Type I seesaw”: $m_I = M_e(M_d - M_e)^{-1}M_e - 6\xi M_e + 9\xi^2(M_d - M_e)$
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“Type II seesaw”: $m_{II} = M_d - M_e$
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**OUR CONVENTION:**  $\hat{M}_j = U_j^T M_j U_j$ ,  $j = u, d, e, n$

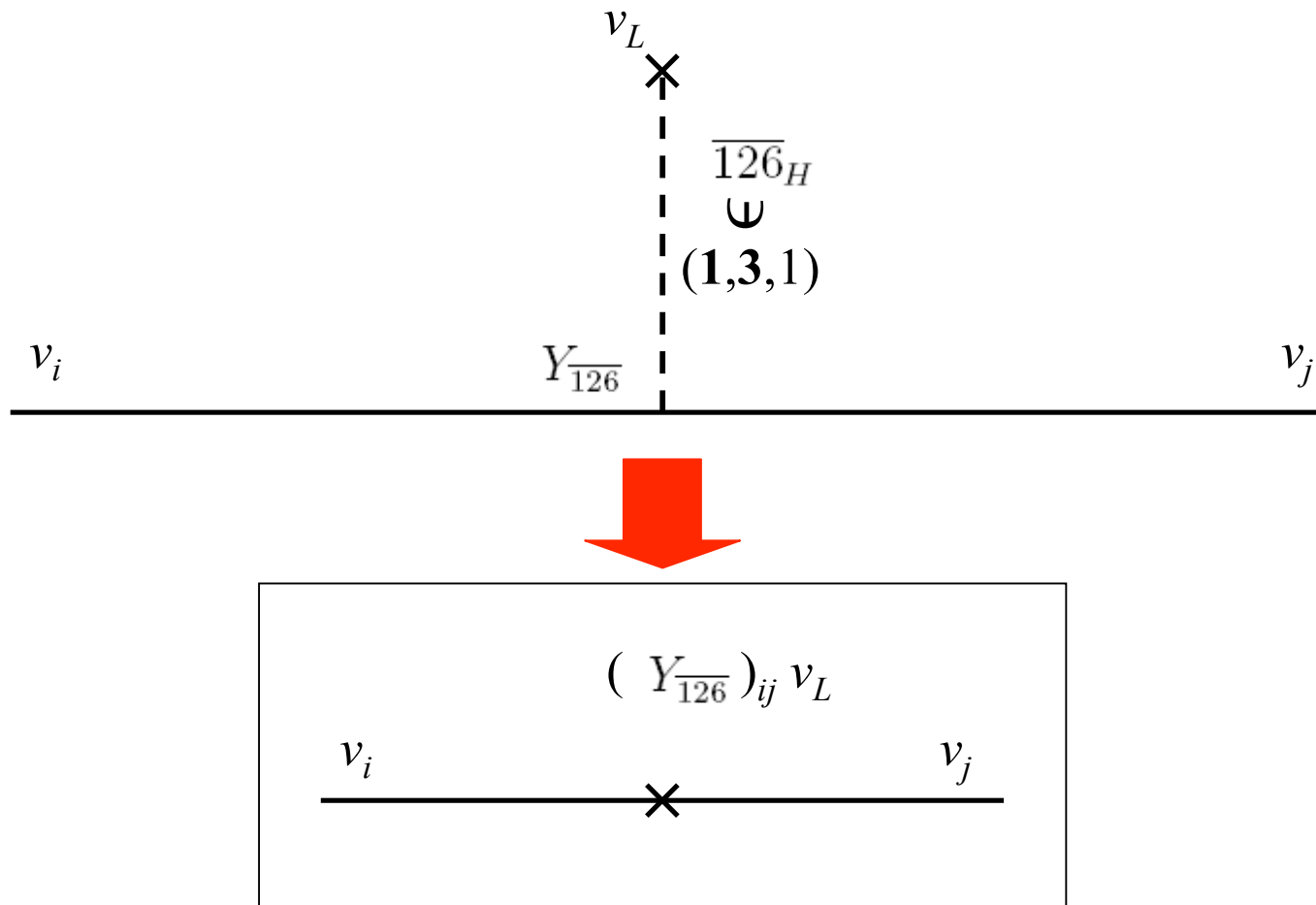
\*B. Bajc, A. Melfo, G. Senjanović and F. Vissani, hep-ph/0511352.

# NEUTRINO MASS — “TYPE I SEESAW”



$v_R \equiv$  vacuum expectation value

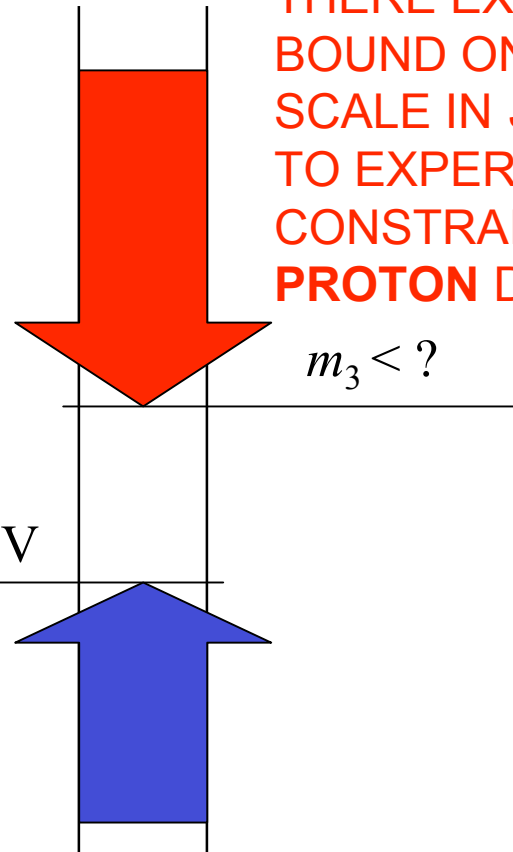
# NEUTRINO MASS — “TYPE II SEESAW”



$v_L \equiv$  vacuum expectation value

# NEUTRINO MASS SCALE PROBLEM IN MINIMAL $SO(10)$

THERE EXISTS AN UPPER  
BOUND ON NEUTRINO MASS  
SCALE IN  $SO(10)$  MODEL DUE  
TO EXPERIMENTAL  
CONSTRAINTS ON THE  $d = 6$   
**PROTON DECAY OPERATORS!**



THERE EXISTS A LOWER  
BOUND ON NEUTRINO MASS  
SCALE DUE TO  
EXPERIMENTAL DATA ON  
**NEUTRINO OSCILLATIONS!**

# “TYPE II SEESAW” NEUTRINO MASSES

$$M_n = F_I m_I + \underline{F_{II} m_{II}}$$

$$m_{II} = M_d - M_e$$

$$F_{II} = \frac{v \sin^2 \beta}{m \cos \beta} \alpha \sqrt{\frac{|\lambda|}{|\eta|}} \frac{N_u^2}{N_d} |f_{II}|$$

\*B. Bajc, A. Melfo, G. Senjanović and F. Vissani, hep-ph/0511352.

## “TYPE II SEESAW” & NEUTRINO MASS

LOWER BOUND:

$$F_{II} \geq 0.2 \times 10^{-9}$$

**$b-\tau$  UNIFICATION**

UPPER BOUND:

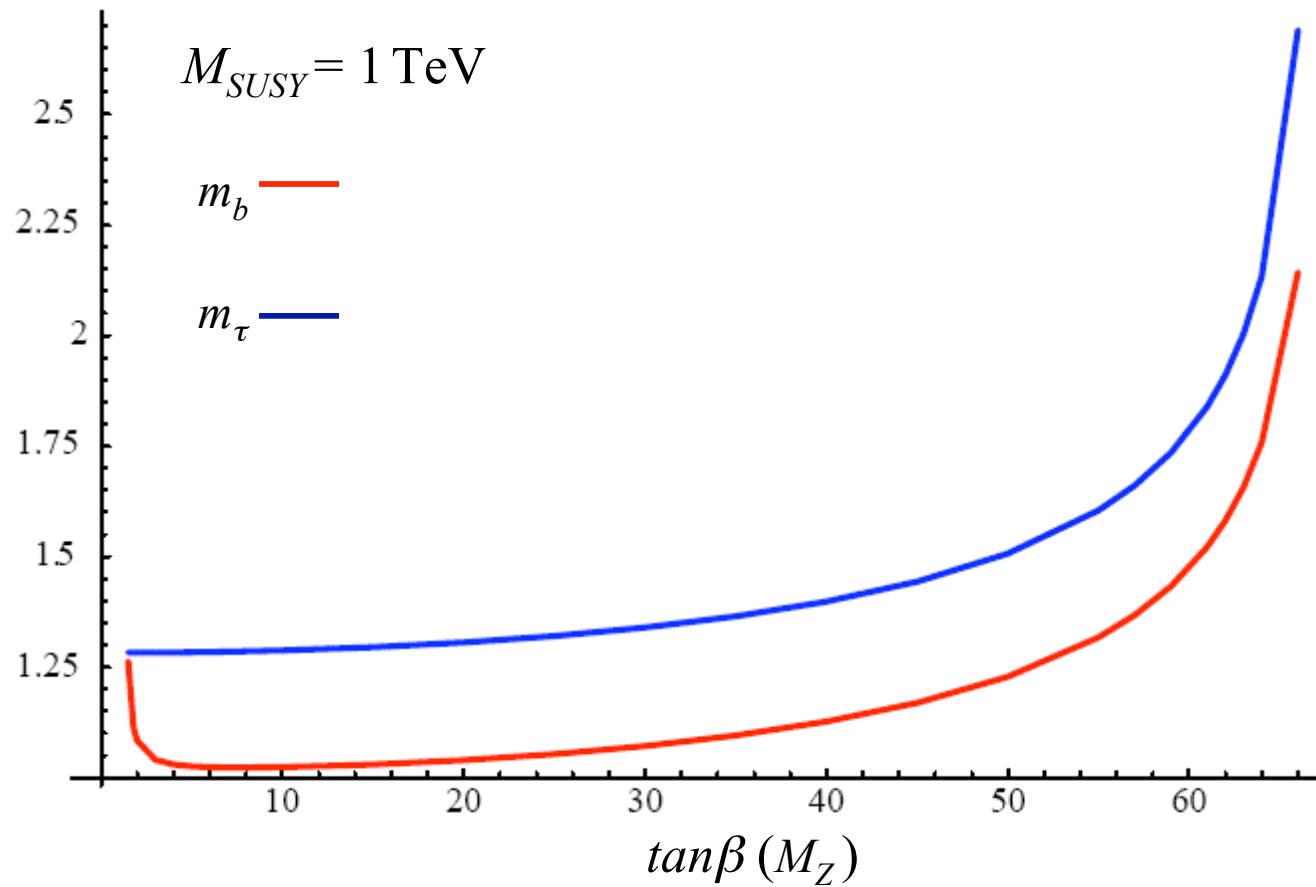
$$\frac{1}{m} < \frac{f(x)}{|\lambda|} \frac{1}{4 \times 10^{15} \text{ GeV}}$$

$$M_{(X,Y)} = m \frac{g_{GUT}}{|\lambda|} \sqrt{4 \left| \frac{1-x-2x^2}{x-1} \right|^2 + 2 \left| \frac{2x(2x^2+x-1)}{(x-1)^2} \right|^2} = m g_{GUT} \frac{f(x)}{|\lambda|}$$

**PROTON DECAY ( $d = 6$ )**

# $b$ - $\tau$ UNIFICATION

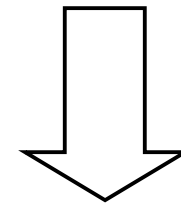
$b$  and  $\tau$  masses at  $M_{GUT}$  in GeV units



$$|m_b - m_\tau| \approx 0.25 \times 10^9 \text{ eV}$$

$$F_{II} \geq 0.2 \times 10^{-9}$$

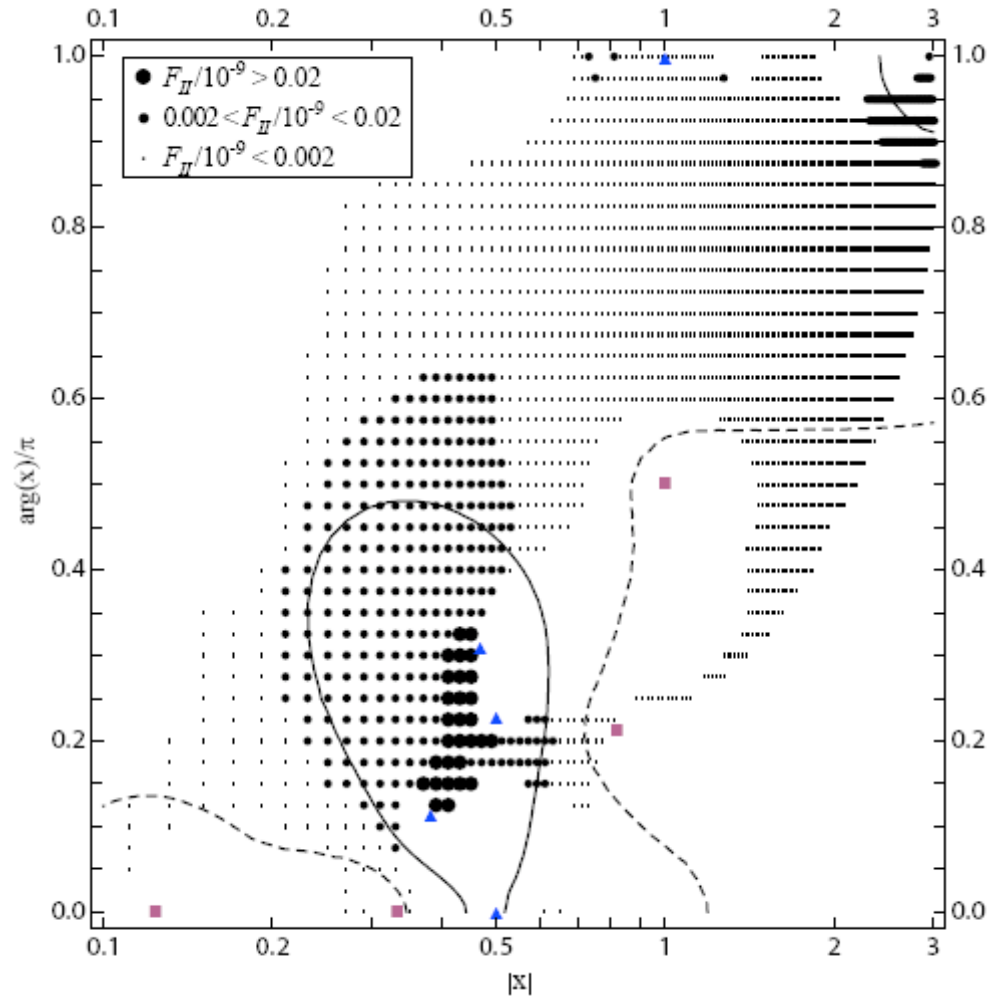
$$F_{II} \times (m_b - m_\tau)$$



$$m_3 \geq 0.05 \text{ eV}$$

# LOW SUSY SCALE & $\nu$ MASS

$$\eta = \lambda = \alpha = \bar{\alpha} = 1 \quad \tan \beta = 50$$





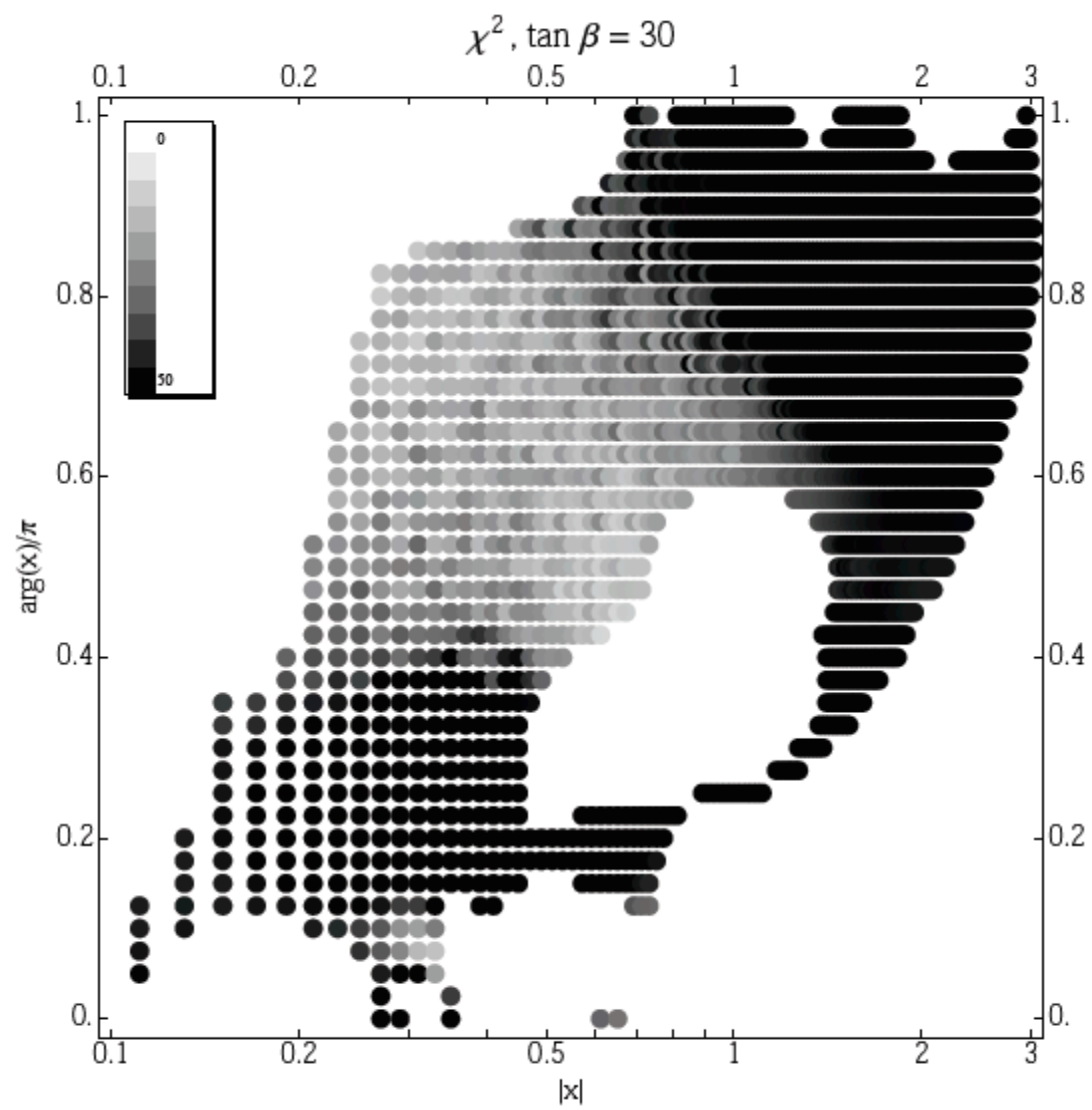
# NUMERICAL PROCEDURE

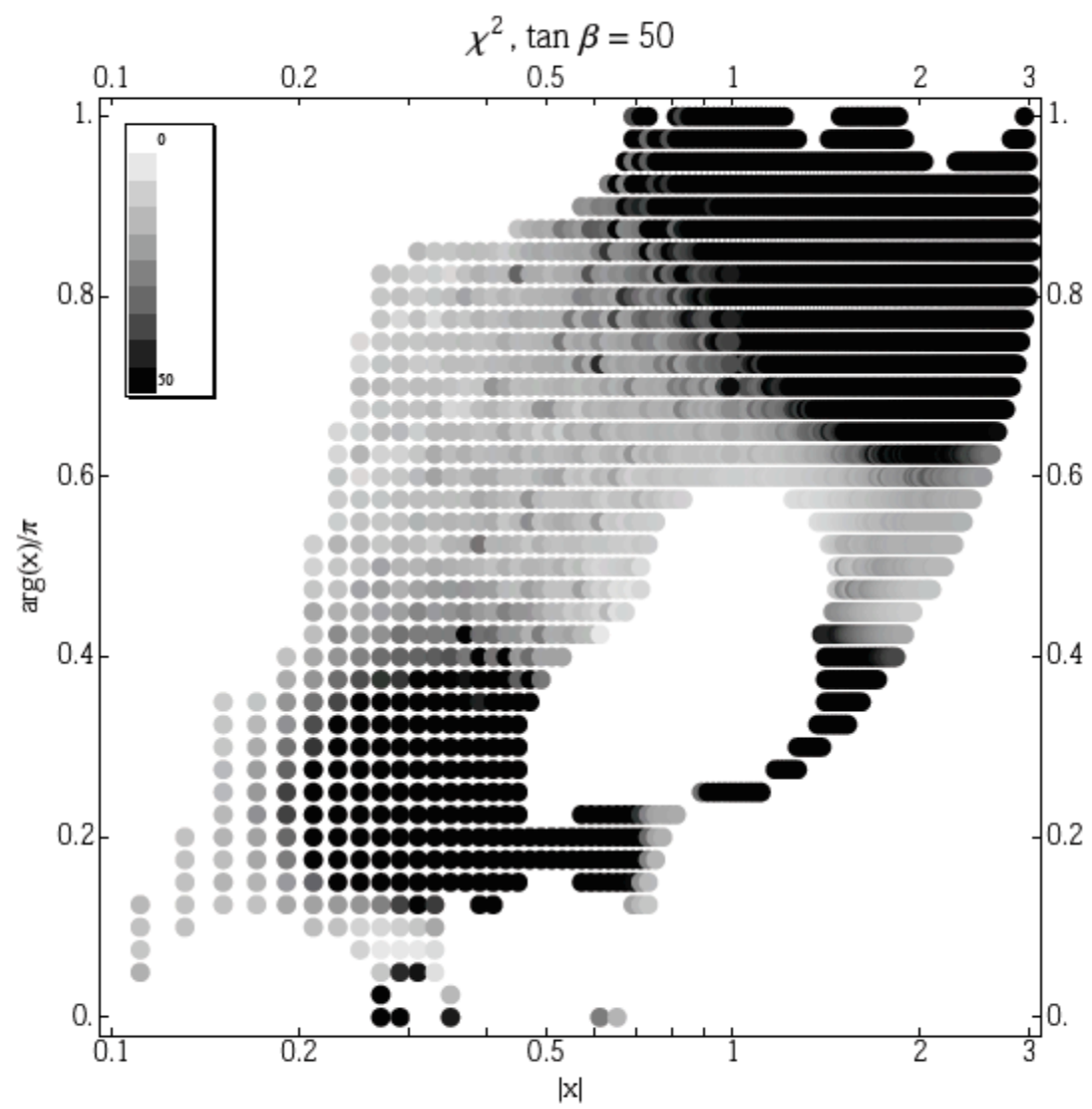
$$\chi_i = \frac{p_i - \tilde{p}_i}{f_i \tilde{p}_i} \quad \longrightarrow \quad \chi^2 = \chi_q^2 + \chi_\ell^2, \quad \chi_q^2 = \sum_{i=1}^{10} \left( \frac{p_i - \tilde{p}_i}{f_i \tilde{p}_i} \right)^2$$

i	1	2	3	4	5	6	7	8	9	10
$p_i$	$m_u$	$m_c$	$m_t$	$m_d$	$m_s$	$m_b$	$s_{12}^{CKM}$	$s_{23}^{CKM}$	$s_{13}^{CKM}$	$\delta^{CKM}$
$f_i$ (%)	30	10	10	35	35	10	7	4	1	20

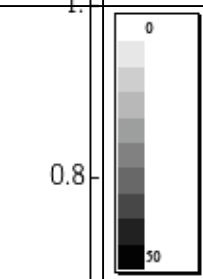
$$\begin{aligned} \chi_\ell^2 = & \left( \frac{m_2/m_3 - 0.18}{0.02} \right)^2 + \left( \frac{s_{12}^{PMNS} - 0.5}{0.15 \times 0.5} \right)^2 \\ & + \left( \frac{s_{23}^{PMNS} - 0.6}{0.1 \times 0.6} \right)^2 + \left( \frac{s_{13}^{PMNS} - 0.17}{0.2 \times 0.17} \right)^2 \Theta (s_{13}^{PMNS} - 0.17) \end{aligned}$$

Notation:  $s_{12}^{CKM} = \sin(\theta_{12}^{CKM}) \dots$



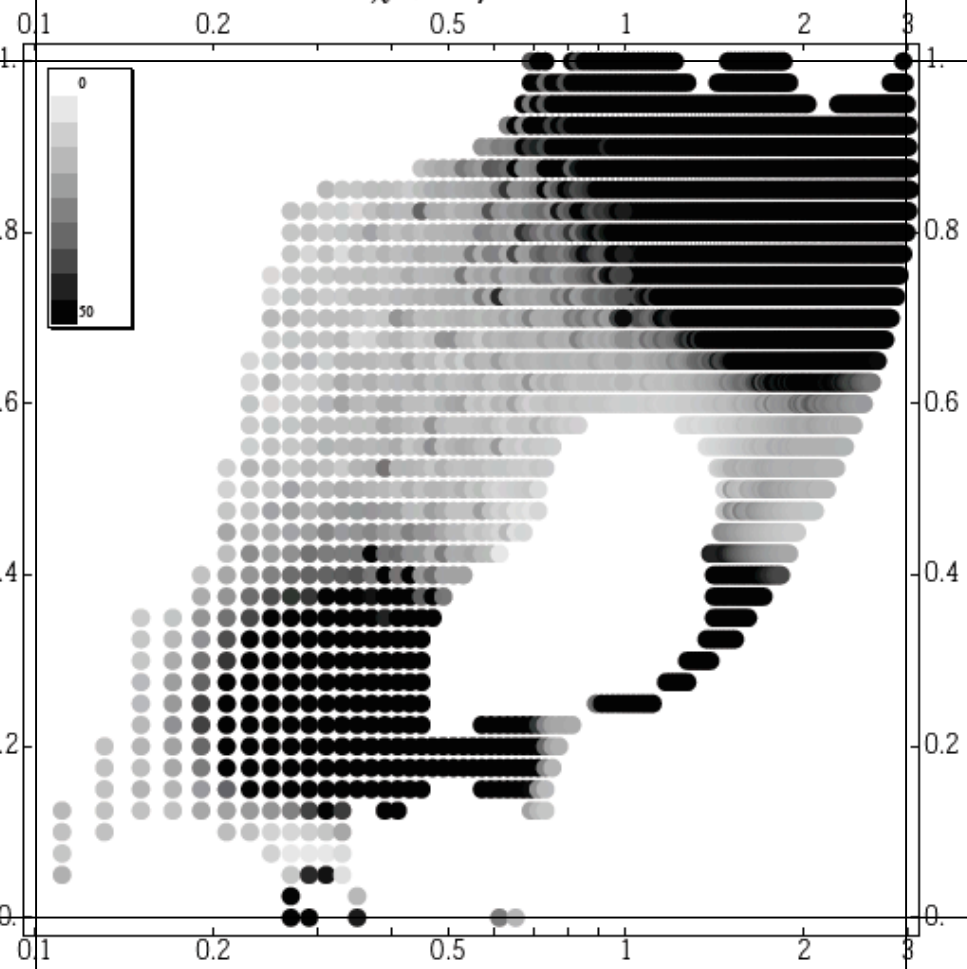


$\chi^2, \tan \beta = 50$

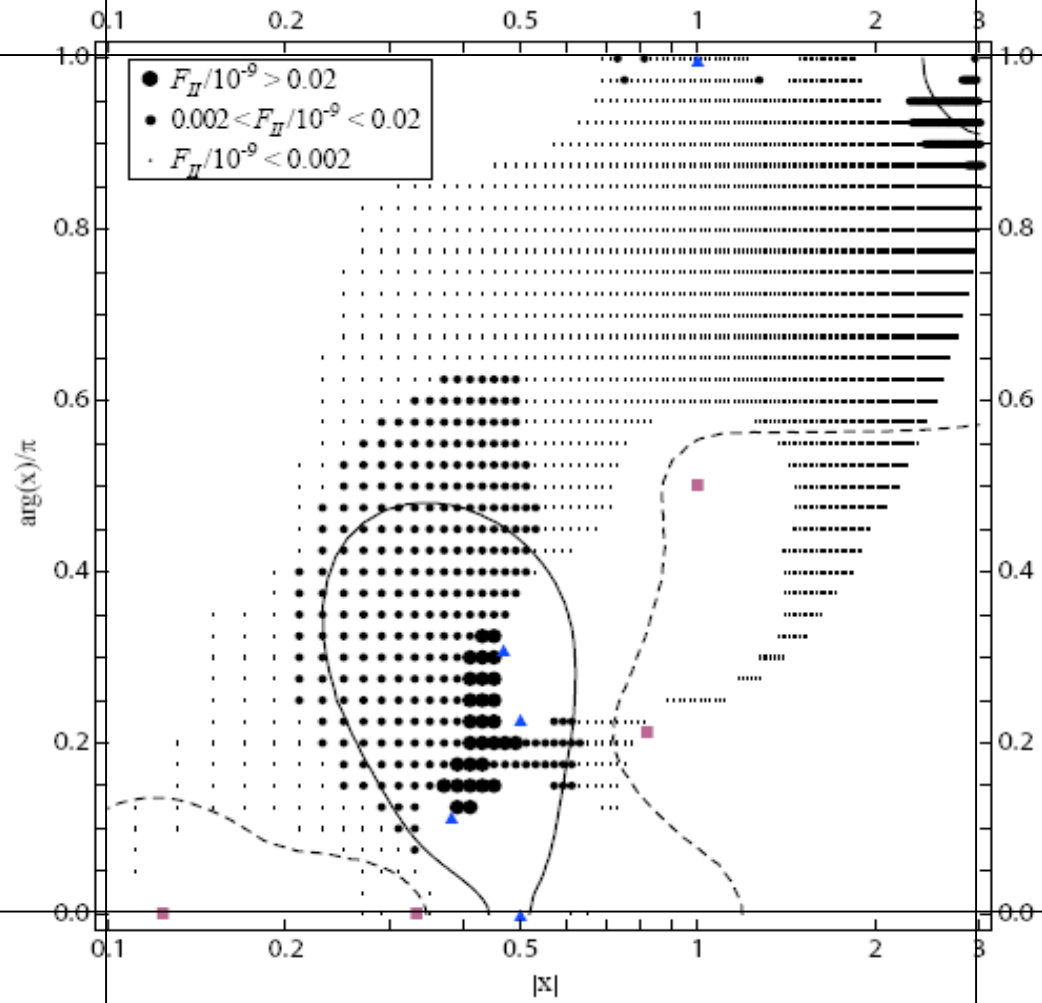


$\arg(x)/\pi$

$|x|$



$$\eta = \lambda = \alpha = \bar{\alpha} = 1 \quad \tan \beta = 50$$



## IMPORTANT FINDINGS...

$$m: \quad 4 \times 10^{14} \text{ GeV} \quad \text{---} \quad 4 \times 10^{17} \text{ GeV}$$

$$\tan \beta = 30$$

	$m_u$	$m_c$	$m_t$	$m_d$	$m_s$	$m_b$	$m_e$	$m_\mu$	$m_\tau$
min	.00053	.210	86.7	.0012	.0213	1.17	.000341	.0720	1.28
max	.00060	.237	93.2	.0014	.0240	1.29	.000367	.0775	1.37
MSSM	.00050	.198	75.7	.0011	.0202	1.07	.000357	.0754	1.34

$$(0.05 \text{ eV}/m_3) \geq 57 \quad \chi^2/\text{d.o.f.} < 10/13$$

NOTE: Charged fermion masses are given in GeV units.

# LOW SUSY SCALE IN MINIMAL $SO(10)$ – STATUS

**According to our findings minimal  $SO(10)$  model with low SUSY scale cannot offer satisfactory description of known fermion masses.**

# “SPLIT SUSY”\* & NEUTRINO MASS

## RELEVANT STEPS:

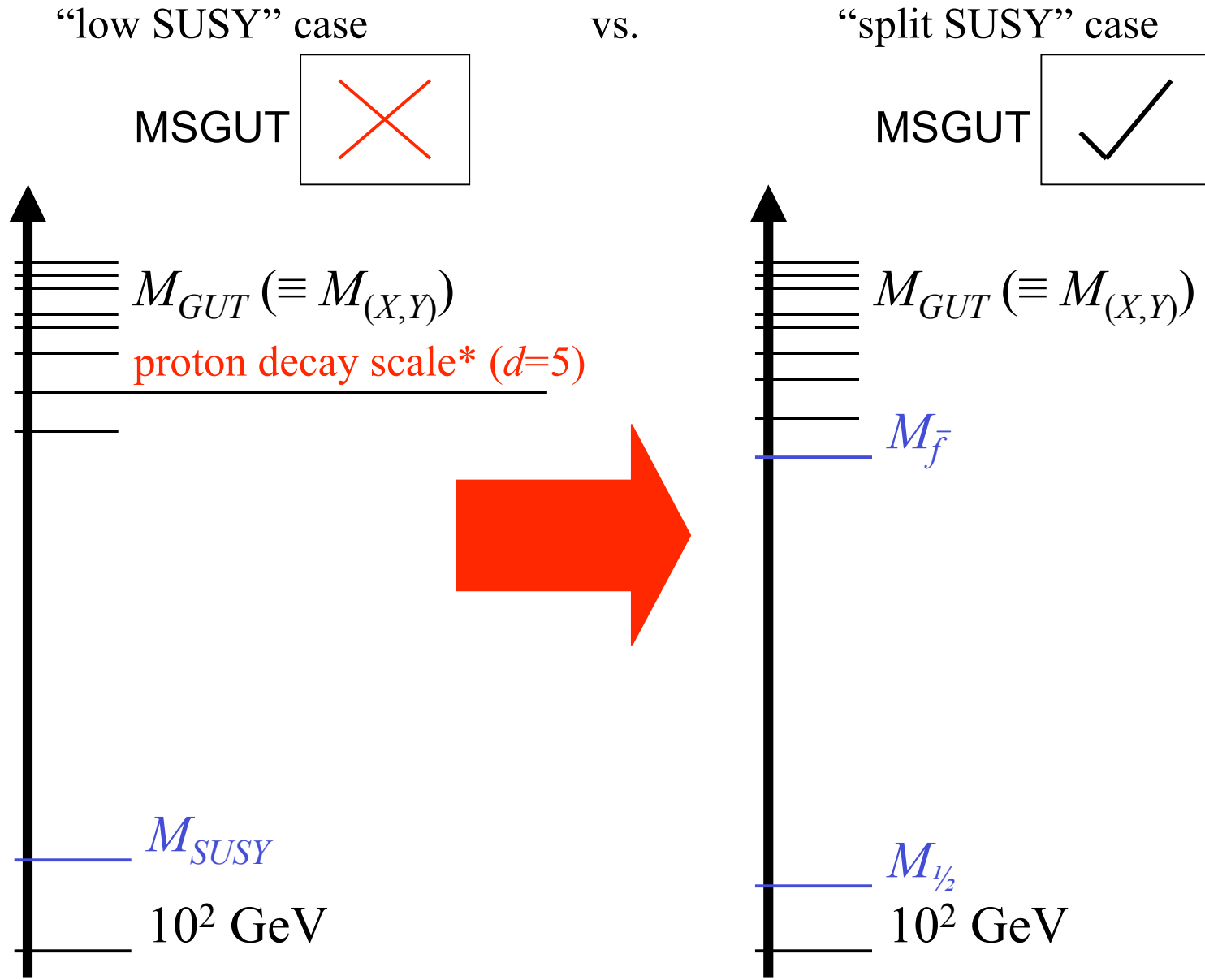
*i)* Find large enough  $F_H$  to describe neutrino mass scale.  
(This step completely determines  $x$ ,  $m$  ( $M_{GUT}$ ),  $\alpha$ ,  $\bar{\alpha}$ ,  $\lambda$  i  $\eta$ .)

*ii)* Check whether the gauge coupling constants unify by solving the relevant renormalization group equations for three unknowns —  $g_{GUT}$ ,  $M_{\bar{f}}$  and  $M_{1/2}$ .

*iii)* Propagate values of quark and charged lepton masses, and CKM parameters to  $M_{GUT}$  for given  $\tan\beta$  and minimize  $\chi^2$ , which also includes neutrino masses ( $m_2/m_3$ ) and the so-called solar and atmospheric angles.

\*N. Arkani-Hamed and S. Dimopoulos, hep-ph/0405159; G. F. Giudice and A. Romanino, hep-ph/0406088.





\*T. Fukuyama et al, hep-ph/0406068; B. Dutta et al, 0712.1206 [hep-ph].

# NUMERICAL RESULTS

observable	data at $M_Z$	FIT	RGE	pull
$m_e$ (MeV)	0.4866613	0.42279	0.42279	
$m_\mu$ (GeV)	0.10273	0.08925	0.08925	
$m_\tau$ (GeV)	1.746	1.534	1.534	
$m_u$ (MeV)	1.6	0.56	0.54	+0.069
$m_c$ (GeV)	0.628	0.218	0.214	+0.182
$m_t$ (GeV)	171.5	72.0	67.4	+0.675
$m_d$ (MeV)	3.5	0.24	1.2	-2.310
$m_s$ (MeV)	62	27.2	21.8	+0.696
$m_b$ (GeV)	2.89	0.917	0.910	+0.079
$s_{12}^{CKM}$	0.2272	0.2423	0.2272	+0.951
$s_{23}^{CKM}$	0.0422	0.0478	0.0474	+0.199
$s_{13}^{CKM}$	0.00399	0.00447	0.00448	-0.015
$\delta^{CKM}$	0.995	1.149	0.995	+0.598
$s_{12}^{PMNS}$		0.42	0.55	-1.091
$s_{23}^{PMNS}$		0.55	0.69	-0.764
$s_{13}^{PMNS}$		0.103		
$m_2/m_3$		0.178	0.180	-0.104

# NUMERICAL RESULTS

Table 2: Input (RGE) and output (FIT) parameters of the numerical fit at the GUT scale with  $\chi^2/\text{d.o.f.} = 9.6/13$ ,  $m_3 = 0.049 \text{ eV}$ ,  $m_2 = 0.0087 \text{ eV}$  and  $m_1 = 0.0012 \text{ eV}$ . Unification takes place for  $|x| = 0.109$ ,  $\text{arg}(x)/\pi = 0.52$ ,  $\alpha = 1.26788$ ,  $\bar{\alpha} = 7$ ,  $\eta = 6.54112$ ,  $\lambda = 0.03$ ,  $M_{1/2} = 1.5 \times 10^5 \text{ GeV}$ ,  $M_{\tilde{f}} = 9.0 \times 10^{13} \text{ GeV}$ ,  $M_{GUT} = 5.8 \times 10^{15} \text{ GeV}$  and  $g_{GUT} = 1.3$  ( $m = 6.5 \times 10^{13} \text{ GeV}$ ). The input data at  $M_Z$  are also given in the second column.

## PROTON DECAY ( $d = 6$ )

THEORY: ( $p \rightarrow \pi^0 e^+$ ):

$$\Gamma = \frac{m_p}{16\pi f_\pi^2} A_L^2 |\alpha|^2 (1 + D + F)^2 \left[ A_{SR}^2 \left| k_1^2 (W_e^\dagger)^{11} + k_2^2 (V_q^\dagger)^{11} (V_q W_e^\dagger)^{11} \right|^2 + A_{SL}^2 k_1^4 \left| (W_e)^{11} + (V_q)^{11} (W_e V_q^\dagger)^{11} \right|^2 \right],$$

$$W_e = \begin{pmatrix} -0.9989 + 0.0260 i & -0.0288 + 0.0227 i & +0.0136 - 0.0023 i \\ -0.0132 - 0.0348 i & +0.8338 + 0.5478 i & -0.0559 - 0.0156 i \\ -0.0096 - 0.0080 i & -0.0323 - 0.0486 i & -0.7749 - 0.6293 i \end{pmatrix}$$

$$V_q = \begin{pmatrix} +0.8936 + 0.3778 i & -0.2298 - 0.0771 i & -0.0011 + 0.0043 i \\ -0.2386 - 0.0412 i & -0.9648 - 0.0909 i & -0.0438 + 0.0191 i \\ +0.0106 + 0.0002 i & +0.0440 + 0.0159 i & -0.9828 + 0.1785 i \end{pmatrix}$$

$$A_{SL} = 2.0, \quad A_{SR} = 2.2$$

# PROTON DECAY\*

$p$ decay modes	Partial mean life ( $10^{33}$ years)		Fraction ( $\Gamma_i/\Gamma$ )
	MODEL	Lifetime bounds:	
$p \rightarrow \pi^0 e^+$	1.6	$> 1.6$	44.3 %
$p \rightarrow \pi^0 \mu^+$	70	$> .473$	1.0 %
$p \rightarrow K^0 e^+$	442	$> .150$	0.2 %
$p \rightarrow K^0 \mu^+$	15	$> .120$	4.7 %
$p \rightarrow \eta e^+$	238	$> .313$	0.3 %
$p \rightarrow \eta \mu^+$	238	$> .313$	0.5 %
$p \rightarrow \pi^0 \bar{\nu}$	1.5	$> .025$	49.0 %

All relevant phases are determined through numerical procedure!

For example,  $\delta^{PMNS} = 6.2!$

All relevant masses, for example,  $M_{(X,Y)}$ , are fixed through unification of gauge couplings  $\alpha_i!$

\*P. Fileviez Pérez, hep-ph/0403286.

## CONCLUSIONS

- **According to our findings minimal  $SO(10)$  model with low SUSY scale cannot describe known fermion masses in satisfactory fashion.**
- **“Split SUSY” scenario not only provides correct neutrino mass scale but it also describes all masses and mixing parameters correctly.**
- **Model predicts *i*) proton decay lifetime that is well within the reach of the next generation of experiments and *ii*) experimentally accessible value for  $s_{13}^{PMNS}$ .**

# CONCLUSIONS

**Minimal  $SO(10)$  is still viable (and predictive)  
candidate of Grand Unification**